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lecture cource on

# Astroparticle physics

15.09.2009 - 15.12.2009











- ▶ 6.1 Double  $\beta$ -decay Decay process
- ▶ 6.2 Double  $\beta$ -decay Motivation
- ▶ 6.3 Double β-decay Candidate isotopes
- ▶ 6.4 Double  $\beta$ -decay Experimental  $\nu$ -mass determination
- ▶ 6.5 Double  $\beta$ -decay Performed experiments
- ▶ 6.6 Double  $\beta$ -decay Future experiments
- ▶ 6.7 Double  $\beta$ -decay Enrichment

#### Decay processes – $\beta$ and $\beta\beta$ decays

### ▶ Normal $\beta$ -decay

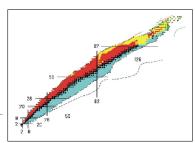
$$\beta^{-}: n \longrightarrow p + e^{-} + \bar{\nu}_{e}$$

$$(Z, A) \longrightarrow (Z + 1, A) + e^{-} + \bar{\nu}_{e}$$

$$\triangleright \beta^+ : p \longrightarrow n + e^+ + \nu_e$$

$$EC: p + e^+ \longrightarrow n + \nu_{e^-}$$

$$(Z, A) \longrightarrow (Z - 1, A) + e^+ + \nu_{e^-}$$



#### ▶ Double $\beta$ -decay

(1) with neutrino 
$$(2\beta 2\nu)$$
  
 $n_{1,2} \longrightarrow p_{1,2} + e_{1,2}^- + \bar{\nu}_{e1,e2}$ 

$$(Z,A) \longrightarrow (Z+2,A) + e_1^- + e_2^- + \bar{\nu}_{e_1}^- + \bar{\nu}_{e_2}^-$$

• (2) neutrinoless decay 
$$(2\beta 0\nu)$$

$$n_{1,2} \longrightarrow p_{1,2} + e_{1,2}^- + \bar{\nu}_{e1,e2}$$
  
 $(Z,A) \longrightarrow (Z+2,A) + e_1^- + e_2^-$ 

Decay processes - semiempirical mass formula

$$\qquad \qquad M(Z,A) = Zm(^{1}H) + Nm_{n} - B(Z,A)/c^{2}$$

► 
$$B(Z, A) = a_v A - a_s A^{2/3} - a_c Z(Z - 1) A^{-1/3}$$
  
 $-a_{sym} (A - 2Z)^2 / A + \delta_p$ 

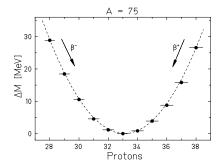
- $a_v = \text{volume term} = 15.5 \text{ MeV}$
- $a_s = \text{surface term} = 16.8 \text{ MeV}$
- $a_c = \text{Coulomb term} = 0.72 \text{ MeV}$
- $a_{sym} = \text{symmetry term} = 23 \text{ MeV}$

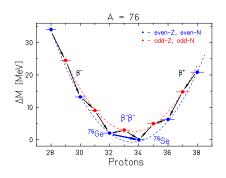
$$\delta_{\rho} = \text{pairing energy} = \begin{array}{c} +a_{\rho}A^{-3/4} & Z, N \text{ even} & (A \text{ even}) \\ 0 & Z \text{ or } N \text{ odd}, A \text{ odd} \\ -a_{\rho}A^{-3/4} & Z, N \text{ odd} & (A \text{ even}) \\ \end{array}$$

Decay processes - mass parabola - fixed A

$$\qquad \qquad M(Z,A) = Zm(^{1}H) + Nm_n - B(Z,A)/c^2$$

► 
$$B(Z, A) = a_v A - a_s A^{2/3} - a_c Z(Z - 1) A^{-1/3}$$
  
 $-a_{sym} (A - 2Z)^2 / A + \frac{\delta}{\delta}$ 





Decay processes

#### Then

- ightharpoonup Double eta-decay may exist for even-even isotopes at the bottom of the valley of stability
- $\blacktriangleright$  Approximately 40 candidiate isotopes exist, out of which most are  $2\beta^-$  type

#### Morever

- ▶ Double  $\beta$ -decay is weak interaction process of the second order  $\bowtie$  half lives are long
- Neutrinoless double  $\beta$ -decay violates lepton number conservation (is forbidden in the standard electroweak theory)
- ► First ideas: Majorana, Racah (1937) and Furry (1939)

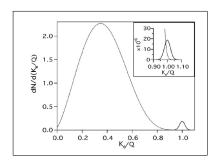
Decay processes

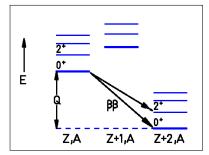
Decay energy (
$$Q_{etaeta}$$
-value)

$$Q_{\beta\beta} = \Delta({}^A_ZX) - \Delta({}^A_{Z+2}X)$$

 $\Delta(\ _Z^{\ A}X\ )$  is the mass defect of isotope (  $_Z^{\ A}X\ )$  (from tables)

For example 
$$Q_{\beta\beta}(^{76}_{32}\text{Ge}) = \Delta(^{76}_{32}\text{Ge}) - \Delta(^{76}_{34}\text{Se})$$
  
= -73212 keV + 75251 keV = 2.039 MeV





Motivation - Neutrinoless case

- Neutrinoless double β-decay is important process for probing the fundamental character of neutrino
- ▶ If it exist, it requires that
  - lepton-number conservation is violated
  - neutrino has non-vanishing mass
  - neutrinos are Majorana particles (neutrino is identical to antineutrino)
- Absolute mass of the neutrino can be determined
- Neutrino experiments show that neutrinos oscillate which means that they should have finite mass
  - the upper current limit (of one flavour) is approximately 40 meV
- ▶ The next-generation  $0\nu2\beta$  experiments should go below that limit
  - ightharpoonup  $t_{1/2}$  comparable to proton decay
  - ▶ material problems does not allow massive detectors ☞

Candidate isotopes –  $Q_{etaeta} \geq 1.7$  MeV

Nuclide	$Q_{etaeta}$	Abun	$T_{1/2}^{2 u}({\sf exp.})$	$T_{1/2}^{0 u}({\sf exp.})$	$\langle m_{ u}  angle$
	[keV]	$\delta$ [%]	[years]	[years]	[eV]
<sup>48</sup> Ca	4272	0.187	$(4.2 \pm 1.2) \times 10^{19}$	$> 9.5 \times 10^{21}$	< 8.3
<sup>76</sup> Ge	2039	7.61	$(1.3 \pm 0.1) \times 10^{21}$	$>1.9\times10^{25}$	< 0.35
				$>1.6\times10^{25}$	< 0.33 - 1.35
<sup>82</sup> Se	2995	8.73	$(9.2 \pm 1.0) \times 10^{19}$	$>2.7\times10^{22}$	< 5
<sup>96</sup> Zr	3350	2.80	$(1.4^{+3.5}_{-0.5})  imes 10^{19}$		
<sup>100</sup> Mo	3034	9.63	$(8.0 \pm 0.6) \times 10^{18}$	$>5.5\times10^{22}$	< 2.1
$^{110}Pd$	2000	11.72			
$^{116}Cd$	2805	7.49	$(3.2 \pm 0.3) \times 10^{19}$	$> 7 \times 10^{22}$	< 2.6
<sup>124</sup> Sn	2287	5.79			
<sup>130</sup> Te	2529	34.08	$(2.7 \pm 0.1) \times 10^{21}$	$>1.4\times10^{23}$	< 1.1 - 2.6
<sup>136</sup> Xe	2468	8.87	$> 8.1 \times 10^{20}$	$>4.4\times10^{23}$	< 1.8 - 5.2
<sup>148</sup> Nd	1929	5.7			
$^{150}Nd$	3367	5.6	$(7.0^{+11.8}_{-0.3})  imes 10^{18}$	$>1.2\times10^{21}$	< 3
<sup>160</sup> Gd	1730	21.86			

Experimental neutrino-mass determination –  $0\nu2\beta$ 

- ▶ Half-life  $(T_{1/2})$  of the decay can be measured
- ► The half-life  $(T_{1/2})$  of  $0\nu 2\beta$ -decay is inversely proportional to the square of the effective mass  $(\langle m_{\nu} \rangle)$  of the Majorana neutrino, the phase-space factor  $(G_{0\nu})$  and the nuclear matrix element (NME)

$$T_{1/2}(0
u)^{-1} = G_{0
u} imes |\mathsf{NME}|^2 imes \langle m_
u 
angle^2$$

- ▶ the (lower limit) of the half-life can be measured
- phase-space factor can be calculated (not difficult)
- the matrix element which involves the nuclear structure information – can also be calculated
  - real can be calculated well enough only to few isotopes
- ▶ The effective neutrino mass,  $U_{ei}$  being the mixing matrix element and  $m_i$  the mass eigenvalue

$$\langle m_{
u} 
angle = \sum_{i} U_{\mathrm{e}i}^{2} m_{i}$$

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Experimental neutrino-mass determination –  $0\nu2\beta$ 

▶ The lower limit of the half-life  $(T_{1/2})$  is obtained from equation

$$T_{1/2} \sim \epsilon \cdot \delta \cdot \left[ \frac{M \cdot \tau}{\Delta E \cdot R_{\text{bg}}} \right]^{1/2}$$

#### where

- ightharpoonup is the efficiency of the detector
- lacksquare  $\delta$  is the abundance or enrichment of the isotope
- ightharpoonup au is measuring time [years]
- M is the total amount of the isotope [kg]
- ▶  $\Delta E$  is the energy resolution (FWHM) [keV]
- $R_{\rm bg}$  is the background at the energy window of  $Q_{0\nu2\beta}$  [counts/(yr-kg-keV)]

Experimental neutrino-mass determination - Ideal experiment - 1

- ▶ The detector mass must be large enough to reach the 50-meV limit  $(\sim 1 \text{ ton of isotope})$
- ▶ The  $\beta\beta(0\nu)$  source must be extremely low in radioactive contamination
- ► Although the use of natural isotope will be less costly, the enrichment process provides a good level of purification and also results in a (usually) smaller volume detector
- ► A small detector volume minimises internal background (which scale with the detector volume). It also minimises external background by allowing smaller shields and structures (active detector)
- ► Good energy resolution
- ▶ Ease of operation due to long experiments and remote locations

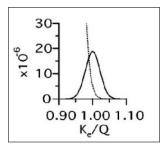
Experimental neutrino-mass determination - Ideal experiment - 2

- ▶ A large  $Q_{\beta\beta}$ -value gives faster  $\beta\beta(0\nu)$  rate and places the region of interest above many potential background activities
- ▶ A relatively slow  $\beta\beta(2\nu)$  rate helps to control the experiment
- ▶ Identifying the daughter in coincidence with the  $\beta\beta$  decay would eliminate most potential background events except  $\beta\beta(2\nu)$
- Good spatial resolution and timing information can help reject background processes
- ► The nuclear theory is better understood in some isotopes than others.
- ► Event reconstruction, providing kinematic data such as opening angle and individual electron energy, can aid in the elimination of background

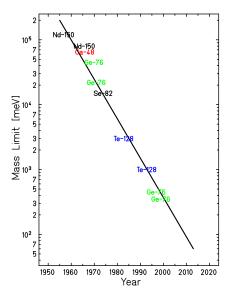
Experimental neutrino-mass determination - Background

- ▶ Double  $\beta$ -decay (both  $2\nu 2\beta$  and  $0\nu 2\beta$ ) experiments are essentially a hard struggling and protection against background radiations
- ► The background can be divided into three groups

  - Cosmic-ray induced activity
  - $2\nu 2\beta$ -decay for  $0\nu 2\beta$ -decay



Experimental neutrino-mass determination – the Moores law for the sensitivity



Performed experiments

The first  $2\beta 0\nu$  experiment : Fireman 1948 [Phys. Rev. 74 (1948) 1238]

Since then  $2\beta 0\nu$  decay studied for 28 isotopes

No  $2\beta 0\nu$  decay observed,  $2\beta 2\nu$  decay observed for 7 isotopes

Isotope	Experimental	Experimental	$\langle m_{\nu} \rangle$	Experiment / Reference
	$T_{1/2}(2\nu)$ [yr]	$T_{1/2}(0\nu)$ [yr]	[eV]	
<sup>48</sup> Ca	$(4.2 \pm 1.2) \times 10^{19}$	> 9.5×10 <sup>21</sup>	< 8.3	You Ke et al., Phys. Lett B265 (1991) 53
76 <sub>Ge</sub>	$(1.3 \pm 0.1) \times 10^{21}$	$> 1.9 \times 10^{25}$	< 0.35	Heidelberg Moscow (1999)
		$> 1.6 \times 10^{25}$	< 0.33 - 1.35	IGEX (1999)
<sup>82</sup> Se	$(9.2 \pm 1.0) \times 10^{19}$	$> 2.7 \times 10^{22}$	< 5	S.R. Elliott et al., PRC46 (1992) 1535
100 <sub>Mo</sub>	$(8.0 \pm 0.6) \times 10^{18}$	$> 5.5 \times 10^{22}$	< 2.1	ELEGANTS (2001)
116 <sub>Gd</sub>	$(3.2 \pm 0.3) \times 10^{19}$	$> 7.0 \times 10^{22}$	< 2.6	F.A. Danevich et al., PRC62 (2000) 04450
<sup>128</sup> Te		$> 7.7 \times 10^{24}$	< 1.1 - 1.5	T. Bernatowicz et al., PRC47 (1993) 806
130 <sub>Te</sub>		$> 1.4 \times 10^{23}$	< 1.1 - 2.6	MIBETA (2000)
136 <sub>Xe</sub>		$> 4.4 \times 10^{23}$	< 1.8 - 5.2	GOTTHARD TUNNEL (1998)

The sensitivity should be improved at least two orders of magnitude

Performed experiments - MIBETA - 1

- ► Material <sup>128,130</sup>Te
- ▶ Detector: Bolometer
  - utilised the low heat capasity of the crystal at low temperature
  - small energy deposit causes significant temperature increment in the crystal
  - good energy resolution
- Consisted of 20 TeO<sub>2</sub> crystals, each of size of 3×3×6 cm<sup>3</sup>, and total mass 6.8 kg
- ► Cooled down to ~10 mK
- Heat and background protection
  - ▶ high-purity electrolytically manufactured copper (2.2 cm)
  - ightharpoonup low-activity lead (Roman lead, < 4 mBq/kg  $^{210}$ Pb) 10 cm
  - ▶ normal lead (16±4 Bq/kg <sup>210</sup>Pb)

Performed experiments - MIBETA - 2

- lacktriangle Energyresolution (FWHM) at 2615 keV :  $\Delta E \sim 9$  keV
- $\delta$ ( $^{130}\mathrm{Te}$ ) = 34%  $\longrightarrow$  enrichment not necessary, still 2 crystals  $\delta$ ( $^{130}\mathrm{Te}$ ) = 93%, and 2 crystals  $\delta$ ( $^{128}\mathrm{Te}$ ) = 95%
- ► Background
  - $R_{bg} \sim 0.5 \text{ counts/(yr-kg-keV)} @ Q_{\beta\beta}(^{130}\text{Te}) = 2529 \text{ keV}$
  - most of it from  $\alpha$ -activity
- Measurement
  - ► Gran Sasso. 3500 mwe
  - ▶ 66995 h  $\times$  crystal  $\longrightarrow$  0.66 kg  $\times$  yr <sup>130</sup>Te :

$$T_{1/2}(0
u) \ge 1.44 \times 10^{23} \text{ yr (}^{130}\text{Te)}$$

$$T_{1/2}(0\nu) \ge 8.6 \times 10^{22} \text{ yr (}^{128}\text{Te)}$$

Performed experiments - Gotthard tunnel

- ► Material <sup>136</sup>Xe
- ► Detector: TPC (Time Projection Chamber)
- ► Active vlume 180 litres of Xe gas
  - $(\delta(^{136}\mathrm{Xe})=62.5\%)$  at pressure of 5 atm 🖾  $9.1{\times}10^{24}$  atomia
- ▶ e<sup>-</sup> tracking → reducing background
- ▶ Energyresolution (FWHM) at  $Q_{\beta\beta}(^{136}{\rm Xe})=2481~{\rm keV}$  :  $\Delta{\rm E}\sim165~{\rm keV}~(6.6~\%)$
- Background
  - $ightharpoonup R_{bg} \sim 0.02 ext{ counts/(yr-kg-keV) @ } Q_{etaeta}(^{136} ext{Xe}) = 2480 ext{ keV}$
  - dominating background at  $2\beta0\nu$  was due to Compton-scattered electrons which were created by natural activity
- ► Measurement
  - Gotthardin highway tunnel (Switzerland?)
  - ▶ 6830 h + 6013 h ( $\sim$ 530 days)  $\longrightarrow$  4.9 kg  $\times$  yr :  $T_{1/2}(0\nu) \ge 4.4 \times 10^{23} \text{ yr (}^{136}\text{Xe)}$

Performed experiments - Heidelberg & Moscow

#### Detector

- ▶ 5 Ge crystals (HPGe),  $\delta(^{76}\text{Ge}) = 86\%$ , mass 10.96 kg (125.5 moles)
- passive and aktive background shielding, PSA
- ▶ Energyresoluution (FWHM) at  $Q_{\beta\beta}(^{76}\text{Xe}) = 2039 \text{ keV}$ :  $\Delta \text{E} \sim 3.9 \text{ keV}$

#### Background

- ▶  $R_{bg} \sim 0.06 \text{ counts/(yr\cdot kg\cdot keV)}$  (with PSA)  $R_{bg} \sim 0.20 \text{ counts/(yr\cdot kg\cdot keV)}$  (without PSA) @  $Q_{\beta\beta}(^{76}\text{Xe}) = 2039 \text{ keV}$
- Measurement
  - ► Heidelberg-Moscow collaboration
  - ► Gran Sasso
  - ▶ 24 kg × yr :  $T_{1/2}(0\nu) \ge 1.6 \times 10^{25} \text{ yr (}^{76}\text{Ge)}$

Performed experiments - IGEX

- ▶ IGEX International Germanium EXperiment
- Detector
  - ▶ 2 kg enriched <sup>76</sup>Ge,  $\delta(^{76}\text{Ge}) = 88\%$  (HPGe)
  - PSA, plastic scintillation detector for muon veto
  - background shielding consisted of 2.5 tons argeological and 10 tons of 70-year-old low-activity lead
  - ▶ Energyresoluution (FWHM) at  $Q_{\beta\beta}(^{76}Xe) = 2039$  keV:  $\Delta E \sim 4$  keV
- Background
  - ▶  $R_{bg} \sim 0.06 \text{ counts/(yr\cdotkg\cdotkeV)}$  (with PSA) @  $Q_{\beta\beta}(^{76}Xe) = 2039 \text{ keV}$
- Measurement
  - Canfranc Underground Laboratory (Spain)
  - ▶ 8.87 kg × yr (116.75 mol·yr) :  $\mathbb{E}^{\mathsf{T}} T_{1/2}(0\nu) \ge 1.57 \times 10^{25} \text{ yr (}^{76}\text{Ge)}$

Performed experiments - HM & IGEX

### Heidelberg & Moscow experiment versus IGEX

HM and IGEX obtained approximately the same results for the half life

Their interpretation of the background was, however, different

- HM: The dominant contribution was created outside the Ge crystal
- ► IGEX: The most restricting factor was the isotope <sup>68</sup>Ge that was an impurity in Ge crystal

Performed experiments - ELEGANTS

- ► ELEctron Gamma-ray Neutrino TeleScopy
- ► Material <sup>100</sup>Mo
- ► Detector: Spectrometer ELEGANT V
  - three drift chambers for tracking electron paths
  - plastic scintillation detectors for measuring energy and times
  - ▶ NaI(TI) scis around the spectrometer for  $\gamma$ -ray elimination
  - protected by copper and lead
  - two  $^{100}$ Mo sources ( $\delta$ =95 %, 20 mg/cm $^2$ ), total mass 171 g  $^{100}$ Mo  $\Longrightarrow$  10 $^{24}$  atoms
  - good background separation but not that efficient or not that good energy resolution than Ge crystals or bolometers

#### Background

- ▶ the most dominating background component at the  $2\beta0\nu$  energy window was concluded to arise from isotopes <sup>208</sup>TI and <sup>214</sup>Bi which were impurities in the Mo source and other detector materials
- ► Measurement
  - Kamioka, Japan

Future experiments - could start in 10 years

Aim to reach sensitivity couple of magnitude lower in half-life measurement

measure masses below 50 meV

$$T_{1/2} \sim \epsilon \cdot \delta \cdot \left[ \frac{M \cdot \tau}{\Delta E \cdot R_{bg}} \right]^{1/2}$$

- ► Choise of detector and detector material
  - detector efficiency ~100 %
     active detector (<sup>76</sup>Ge, <sup>116</sup>Gd, <sup>130</sup>Te, <sup>136</sup>Xe)
  - ▶ good energyresolution (FWHM) 🖾 at least  $\sim$ 4%  $\times$   $Q_{\beta\beta}$
  - ▶ HM:  $m(^{76}Ge) \sim 10 \text{ kg}$  Is hundreds of kilos enriched material
  - ▶ large  $Q_{\beta\beta}$
- ▶ Background 🖾 as small as possible at the  $Q_{\beta\beta}$  range
- ▶ Measurement time long (~10 years) 

  detector easy to operate

#### Future experiments – Proposed experiments

Experiment	eriment Source Detector description		Sensitivity to $T_{1/2}^{0\nu}$ [years]	Reference
COBRA	130 <sub>Te</sub>	10 kg CdTe semiconductors	1/2 1×10 <sup>24</sup>	K. Zuber, Phys. Lett B519 (2001) 1
DCBA	150 <sub>Nd</sub>	20 kg <sup>enr</sup> Nd layers	$2 \times 10^{25}$	N. Ishihara et al., NIM A443 (2000) 101
NEMO-3	100 <sub>Mo</sub>	between tracking chambers 10 kg of $\beta\beta(0\nu)$ isotopes (7 kg Mo) with tracking	4×10 <sup>24</sup>	X. Sarazin, hep-ex/0006031
CAMEO	<sup>116</sup> Cd	1 ton CdWO <sub>4</sub> crystals in	$\geq 10^{26}$	G. Bellini et al., EPJ C19 (2001) 43
CANDLES	<sup>48</sup> Ca	liquid scintillator several tons of CaF <sub>2</sub> crystals in liquid scintillator	1×10 <sup>26</sup>	T. Kishimoro <i>et al.</i> , Ann. Rep. Osaka Univ. Lab. Nucl. Studies (2000)
CUORE	130 <sub>Te</sub>	750 kg TeO <sub>2</sub> bolometers	2×10 <sup>26</sup>	F.T. Avignone <i>et al.</i> , hep-ex/0201038
EXO	136 <sub>Xe</sub>	1 ton enr Xe TPC (gas or liquid)	8×10 <sup>26</sup>	F.A. Danevich et al., PR C62 (2000) 044501
GEM	76 Ge	1 ton enr Ge diodes in liquid N	$7 \times 10^{27}$	Yu.G. Zdesenko et al., J.Phys. G27 (2001) 212
GENIUS	$^{76}$ Ge	1 ton <sup>76</sup> Ge (enr. 86%) diodes	$1\times10^{28}$	H.V. Klapdor-Kleingrothaus, hep-ph/0103074
GSO	<sup>160</sup> Gd	in liquid N 2 ton Gd <sub>2</sub> SiO <sub>5</sub> crystal scintillator	2×10 <sup>26</sup>	F.A. Danevich et al., NP A694 (2001) 375 S.C. Wang et al., hep-ex/0009014
Majorana	76 <sub>Ge</sub>	0.5 ton 86% segmented enr Ge diodes	$3 \times 10^{27}$	C.E. Aalseth et al., hep-ex/0201021
MOON	<sup>100</sup> Mo	34 ton <sup>nat</sup> Mo sheets between	$1\times10^{27}$	H. Ejiri et al., PRL85 (2000) 2917
Xe	136 <sub>Xe</sub>	1.56 ton of <sup>enr</sup> Xe in liquid sci.	$5 \times 10^{26}$	B. Caccianiga et al., ApP14 (2001) 15
XMASS	136 <sub>Xe</sub>	10 ton of liquid Xe	$3 \times 10^{26}$	S. Moriyama et al., XENON01 Workshop, Toky

Future experiments - Proposed experiments - CUORE

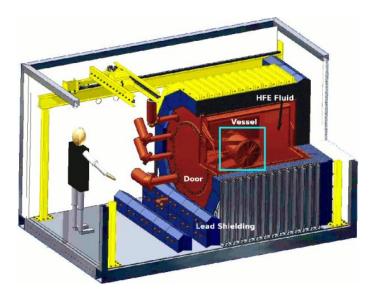
- ► Cryogenic Underground Observatory for Rare Events
- Successor of the MIBETA detector, in Gran Sasso
- ► The detector: Bolometer
  - ▶ 988 TeO<sub>2</sub> crystals with a mass of 750 g each and size of a cube of 5 cm<sup>3</sup>
  - operational temperature  $\sim \! 10$  mK
  - ▶ good energy resolution: 5 - 10 keV (FWHM) at 2.5 MeV
- ► Background protection
  - pure copper and low-activity lead
  - estimated 0.5 0.05 counts/(kg·keV·yr)
- Sensitivity:  $m_{\nu} \leq 20 50 \text{ meV}$
- Prototype CUORICINO, a single CUORE tower is being constructed



Future experiments - Proposed experiments - EXO - 1

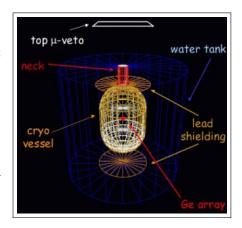
- ▶ The Enriched Xenon Observatory
- ▶ Detector: EXO-200 200 kg prototype experiment currently being installed at WIPP – Waste Isolation Pilot Plant, USA
- ► Liquid xenon TPC
- ▶ Measure  $2\nu2\beta$  mode of 80% enriched <sup>136</sup>Xe  $\Longrightarrow$  limit for  $0\nu2\beta$
- ► Ultrapure materials
- ► EXO a ton-scale experiment <sup>136</sup>Xe
  - under R&D (background elimination)
- ▶ <sup>136</sup>Xe
  - relatively easy to purify the LXe
  - relatively easy to enrich in Russian centrifuges
  - $Q_{\beta\beta} = 2.48$  MeV is quite high
  - ► liquid: scintillation and ionisation
  - background elimination by tagging in the daughter isotope <sup>136</sup>Ba

Future experiments – Proposed experiments – EXO – 2



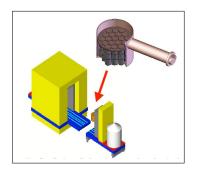
Future experiments - Proposed experiments - GERDA

- GERDA is the successor of the HM and IGEX experiments
- New double  $\beta$ -decay experiment at Gran Sasso, proposed 2004
- ► enriched <sup>76</sup>Ge
- Aim: sensitivity of  $T_{1/2}(0\nu) \geq 1.4^{26}$  yr and  $m_{\nu} \leq 90 300$  keV
- ► Background reduction
  - ► cryostat filled with LN or LAr ⇒ radiopure materials



Future experiments - Proposed experiments - Majorana

- Detector
  - ▶ 120 kg of HPGe crystals,  $\delta(^{76}\text{Ge}) = 86 \%$
- Background protection
  - conventional low-active cryostat, very pure copper  $(< 25 \mu Bq/kg^{226} Ra,$ 9  $\mu$ Bq/kg <sup>228</sup>Th)
  - detectors protected by low-activity lead and copper
  - segmented detectors and PSA
- ► Aim: to achieve background level of  $\sim 1$  counts/(ton·yr) in 4-keV region around  $Q_{\beta\beta}$ if yes  $\Longrightarrow$  scale to 1 ton or larger
- If funding obtained construction started this year



Enrichment of detector materials

Only enrichment plants in Russia are capable, at the moment, to produce hundreds of kilograms of enriched materials

Electro Chemical Plant (ECP), Krasnojarsk

#### Enriched production currently

- ▶  $^{76}$ Ge :  $\sim$ 20 kg/yr
  - with a device never used for Uranium enrichment
- ►  $^{136}$ Xe :  $\sim$ 300 kg/yr

#### With investments

- ▶  $^{76} \text{Ge} : \sim 30 \text{ kg/yr} \longrightarrow \sim 50 \text{ kg/yr} \longrightarrow \sim 100 \text{ kg/yr}$
- ► <sup>136</sup>Xe : ~2 tonnia/yr